

Materials Science & Applications Group



VANDERBILT
UNIVERSITY

Institute of Imaging Science



RADIATION DETECTORS FOR SPACE APPLICATIONS: DETECTOR DESIGN, FABRICATION AND CHARACTERIZATION

Presentation at

*NASA HBCU/MSI Technology Infusion Road Tour
University Capabilities*

September 29, 2016

Emmanuel Rowe, Ph.D.

Materials Science and Applications Group



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- Julia Bodnarik & colleagues, Lunar & Planetary Laboratory, University of Arizona, Tucson, AZ
- Ashley Stowe, Y-12 National Security Complex Oak Ridge, TN
- Nerine Cherepy & Stephen Payne, LLNL, Livermore, CA
- Sebastien Kerisit, Pacific Northwest National Lab, Richland, WA
- Lynn Boatner, ORNL, Oak Ridge, TN
- Kanai Shah and Josh Tower, Radiation Monitoring Devices, Watertown, MA

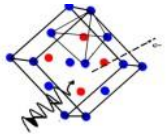
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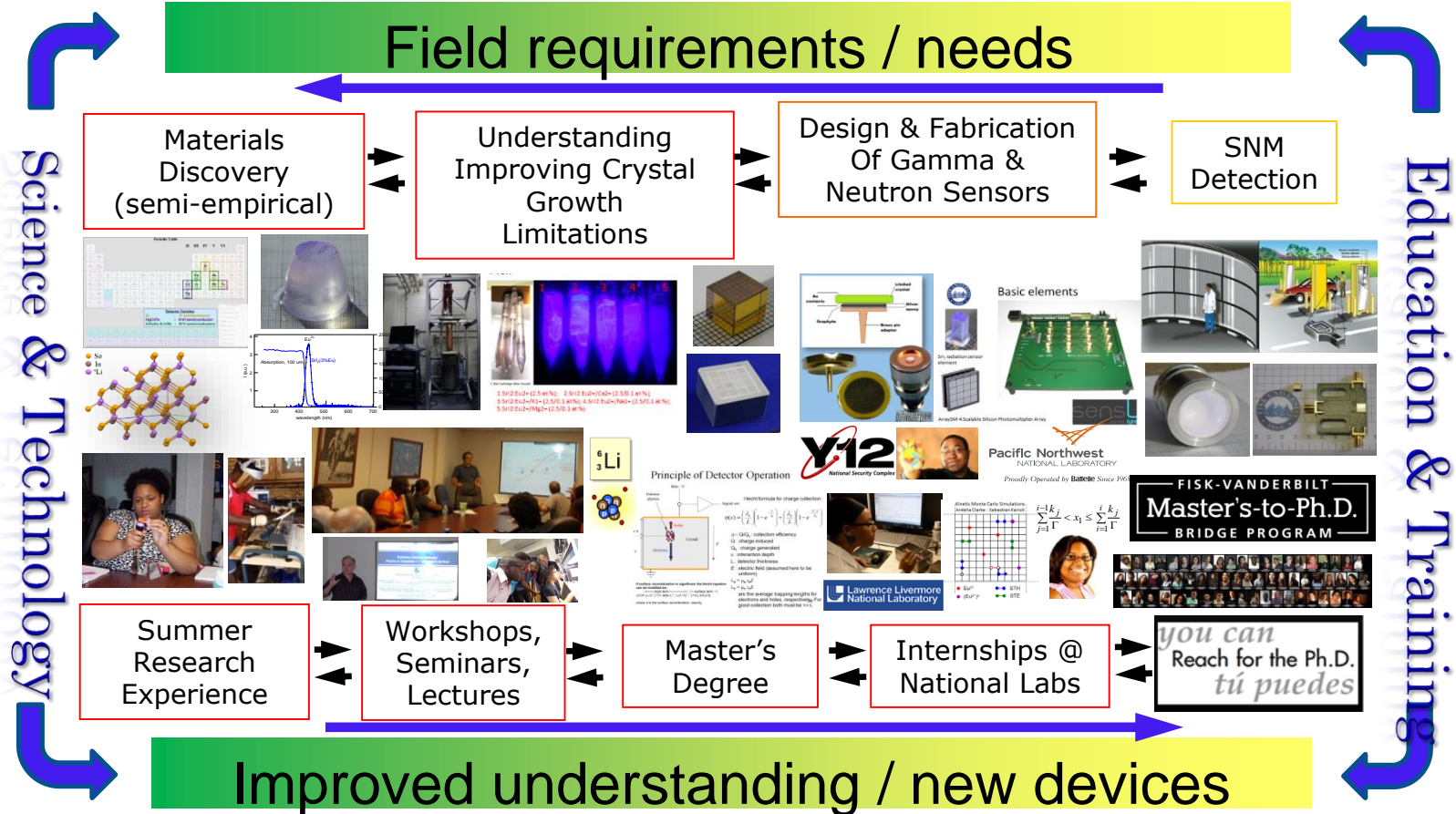
Fisk-Vanderbilt Master's-to-PhD Bridge Program:



Support: NSF, DOE/NNSA, DHS/DNDO, DOD/DTRA,
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Materials Science and Applications Group



Crystal growth is the critical enabling

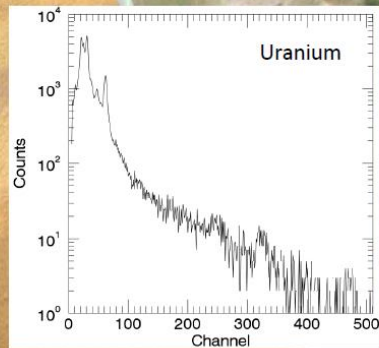
Gamma Ray Spectrometer (GRS)

- Pandora: Unlocking the mysteries of the moons of Mars
- Chemical composition of planetary surfaces provides clues about their origins and evolution

Pandora

Technology development opportunity

1-1.5 u Gamma Ray Spectrometer



Fisk Univ. prototype: (1 u = $10^{-10} \times 10^{-10} \times 10^{-10} \text{ cm}^3$)
 $\text{SrI}_2(\text{Eu})$ Crystal ($15 \times 15 \times 10 \text{ mm}^3$)
with SiPM readout

Tom Prettyman, et al,
2015 SPIE Optics and
Photonics, International
Conference on Hard X-Ray,
Gamma Ray and Neutron
Detector Physics

- Power < 3 W
- Potential for commercial-off-the-shelf components

SrI₂ SCINTILLATORS FOR PLANETARY GAMMA-RAY SPECTROSCOPY

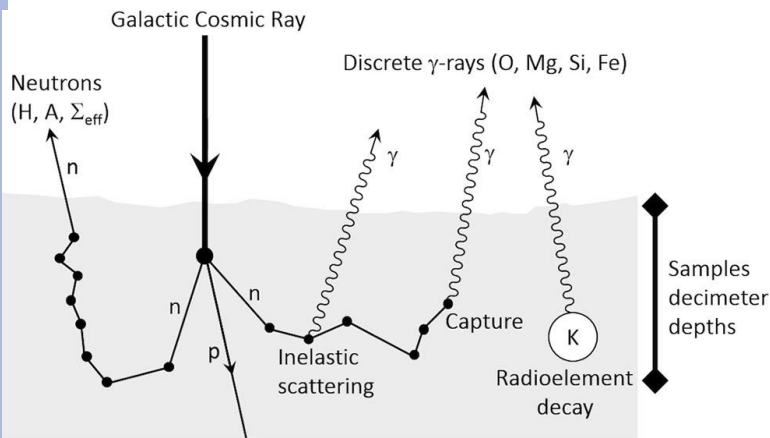


Figure 1. Cartoon of gamma-ray production processes in the surface of planetary bodies. Galactic cosmic rays

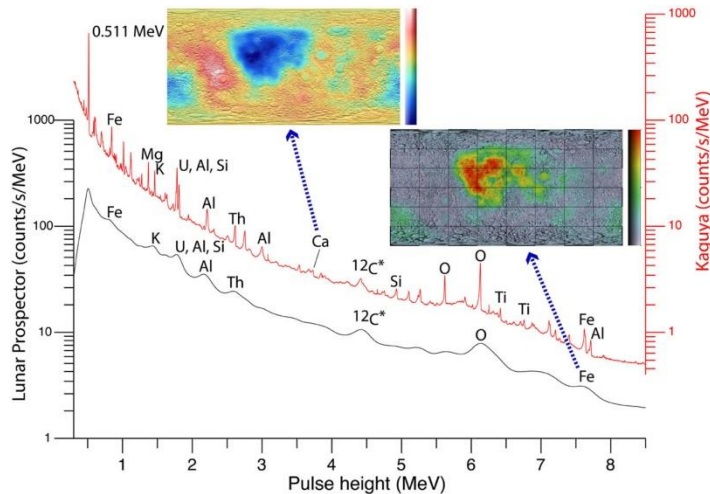


Figure 2. Comparison of the average lunar gamma-ray spectrum, as acquired during the Lunar Prospector (black) and Kaguya (red) missions.

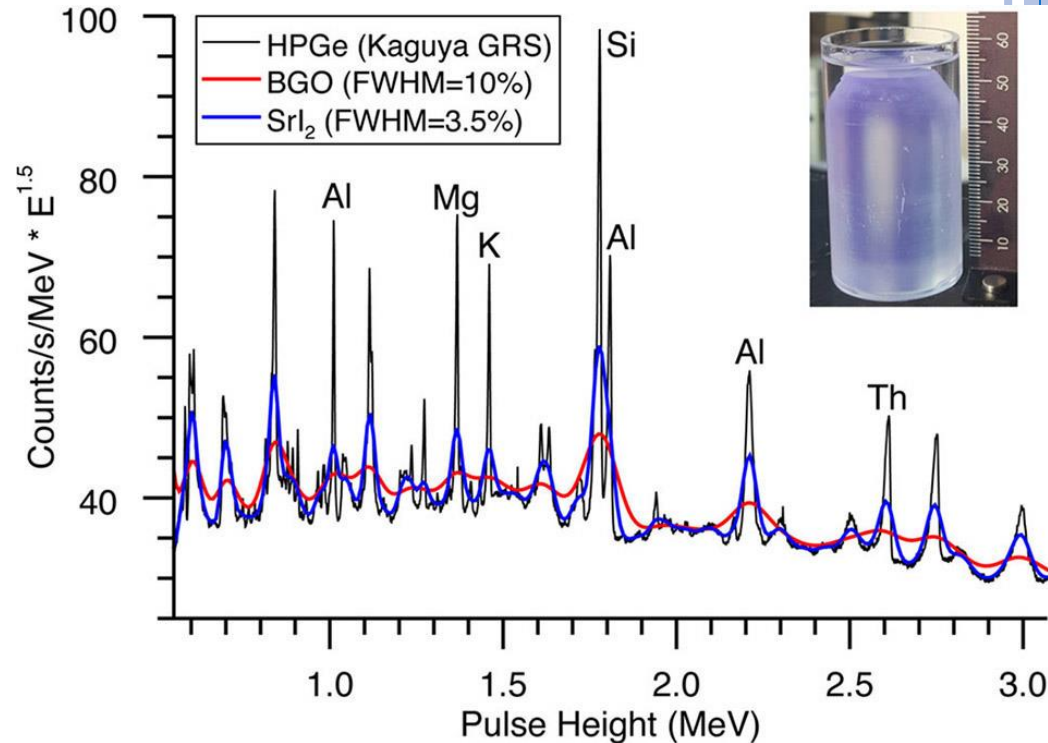
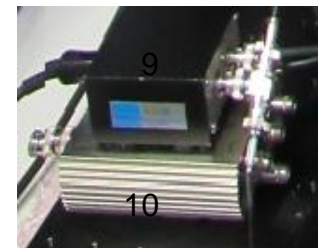
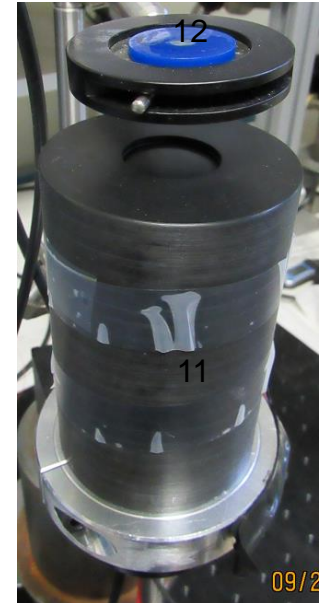
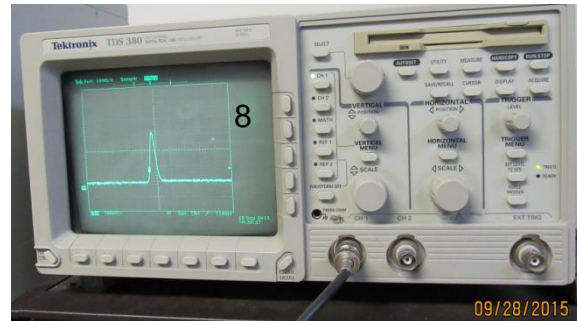
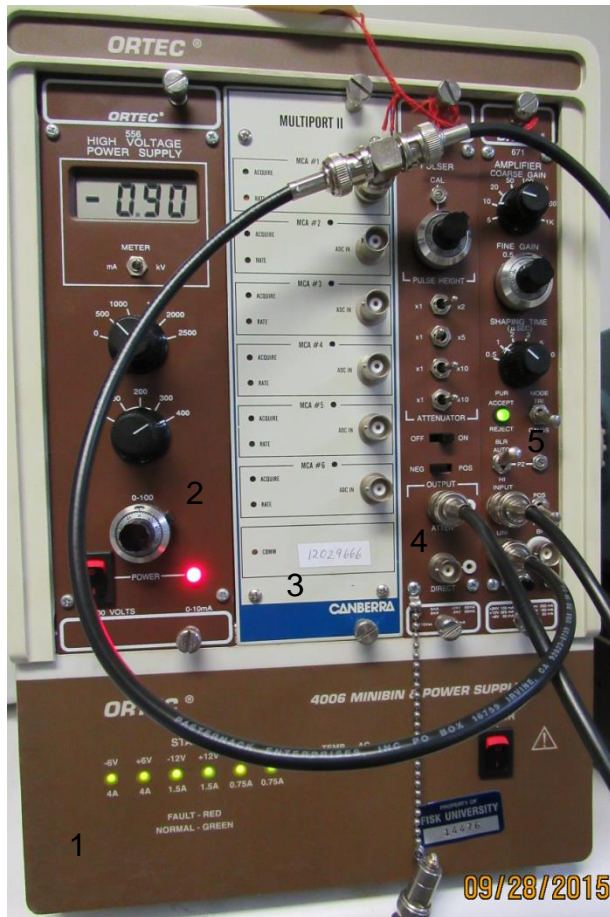


Figure 3. Illustration showing how energy resolution affects the ability to extract chemical information from a lunar gamma-ray spectrum.

Thomas Prettyman, Arnold Burger, Naoyuki Yamashita, James Lambert, Keivan Stassun and Carol Raymond,
23 October 2015, SPIE Newsroom

<http://spie.org/x115974.xml>

BASIC EQUIPMENT USED IN RADIATION DETECTION

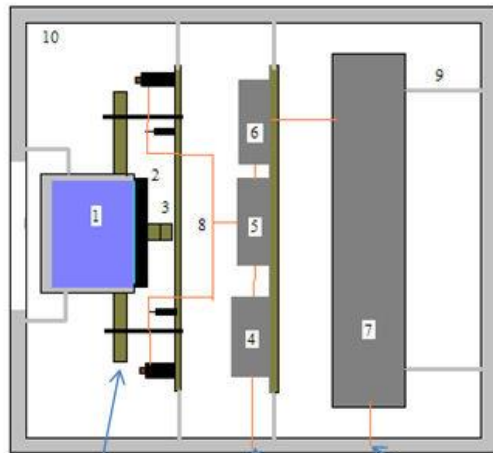


Legend: 1-NIM Minibin, 2-High Voltage Power Supply, 3-Multichannel Analyzer, 4- Pulser, 5- Shaping amplifier
 6-Low Voltage Power Supply, 7-Digital oscilloscope, 8- Analog Oscilloscope, 9-Fast preamplifier C6438
 (for scintillators), 10-Charge sensitive preamplifier (A250CF), 11-PMT(R6231-100), 12-Radiation source

Design



SrI₂/Silicon Photomultiplier-Based Gamma Ray Spectrometer



1. SrI₂ crystal
2. SiPM, SensL ArrayC-30035-16P-PCB
3. Sensor test board, SensL ArrayX-BOB3-16P-V1P1
4. Power supply
5. Preamplifier
6. Shaping amplifier
7. Multi-channel analyzer Kromek K102
8. Wiring
9. Brackets
10. CubeSat chassis and walls assembly



2.4 cm³ SrI₂ scintillator coupled with SensL SiPM

Power supplies
5VDC to
+/- 12 VDC
and -25 to
-30 VDC adj.

Kromek MCA
(multi channel
analyzer)



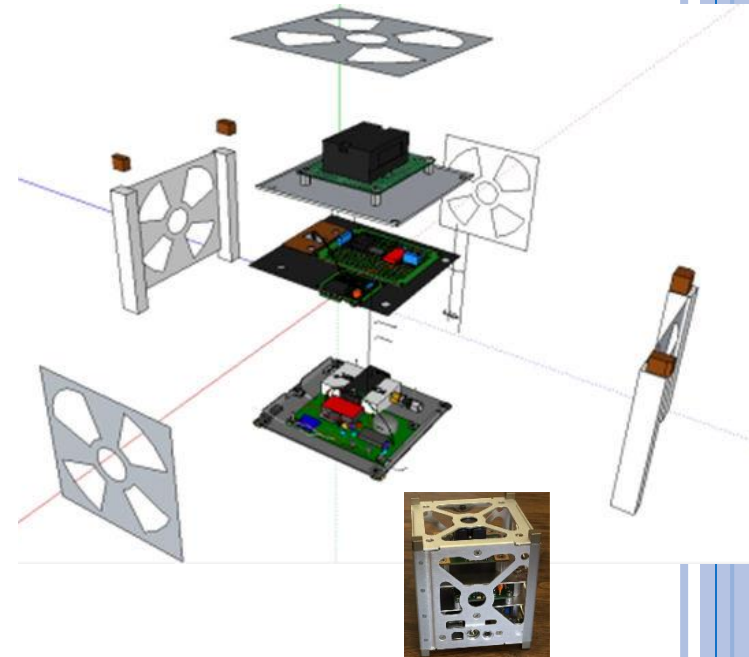
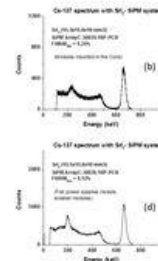
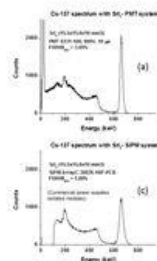
Power supplies and MCA

Cremat
preamplifier
Cr-110

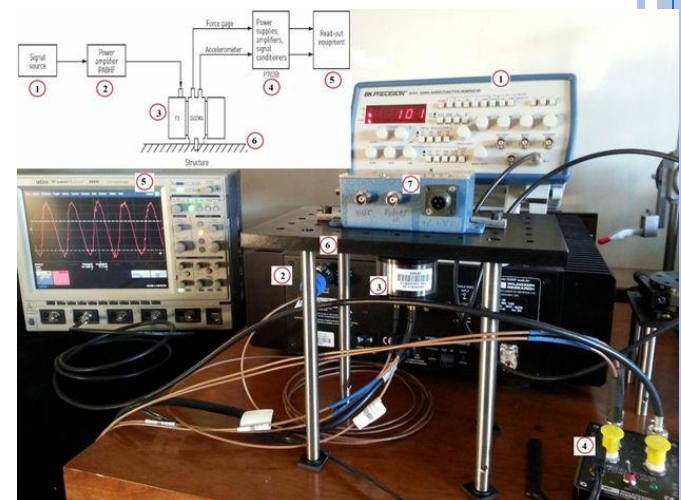
Cremat shaping
amplifier
Cr-200-4µs



Cremat preamplifier and shaping amplifier

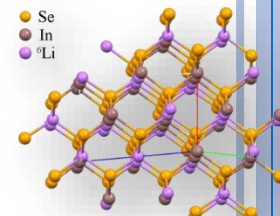


Vibration testing at Fisk



LITHIUM INDIUM DISELENIDE (${}^6\text{LiInSe}_2$)

APPLIED PHYSICS LETTERS 101, 202101 (2012)



Single crystal of LiInSe_2 semiconductor for neutron detector

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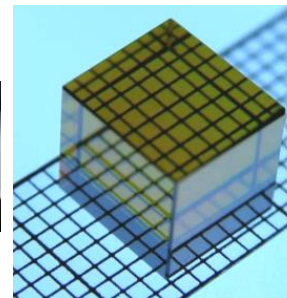
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³Y-12 National Security Complex, Oak Ridge, Tennessee 37830, USA

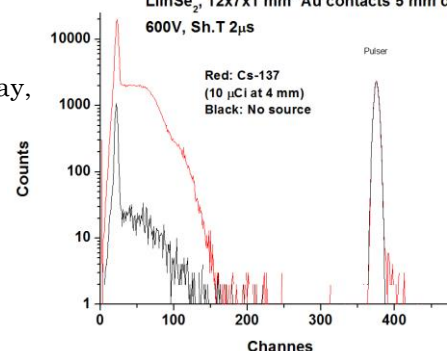
(Received 17 July 2012; accepted 8 October 2012; published online 12 November 2012)

Single crystals of semiconductor-grade lithium indium selenide (LiInSe_2) were grown using the vertical Bridgman method. The orthorhombic structure of the materials was verified using powder x-ray diffraction. The room temperature band gap of the crystal was found to be 2.85 eV using optical absorption measurements. Resistivity of LiInSe_2 , obtained using current-voltage measurements, has semiconducting nature (decreases with increasing temperature) and is in order of $10^{10} \Omega\text{-cm}$. Photoconductivity measurement showed the photocurrent peak at 445 nm. Nuclear radiation devices were fabricated, and alpha particle detection was observed, suggesting that this material could be a candidate for neutron detection applications. © 2012 American Institute of Physics. [<http://dx.doi.org/10.1063/1.4762002>]



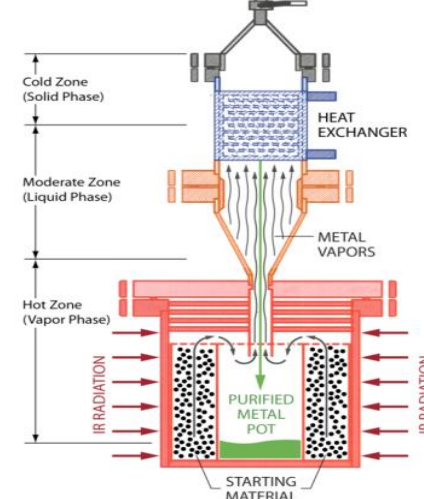
LiInSe_2 Detection response

LiInSe_2 , 12x7x1 mm³ Au contacts 5 mm dia.
600V, Sh.T 2μs

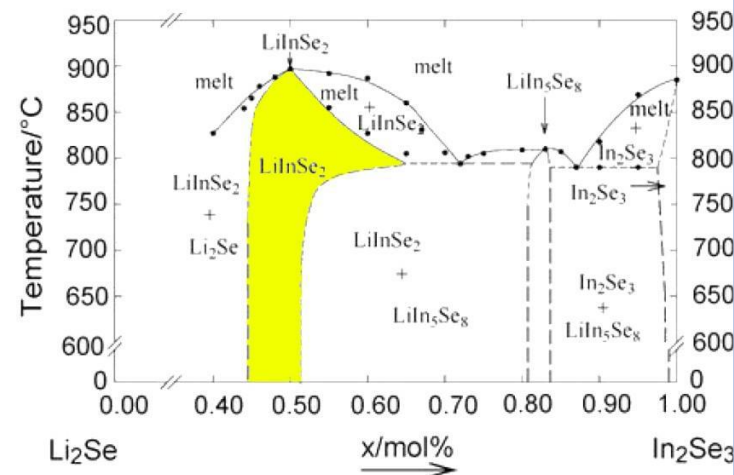
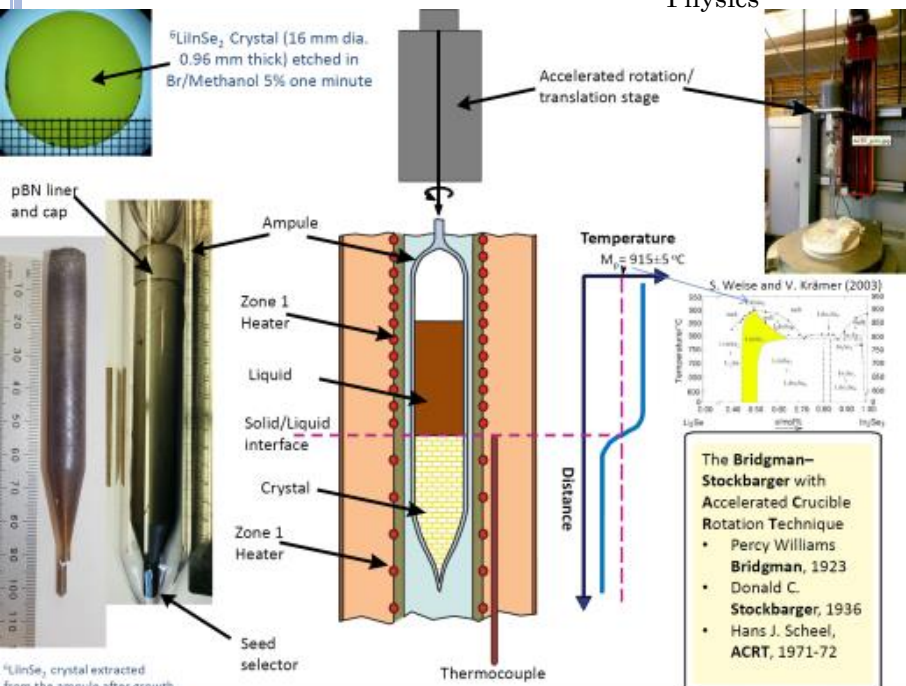


Lithium Distillation - Scale

up
Gas vent/vacuum

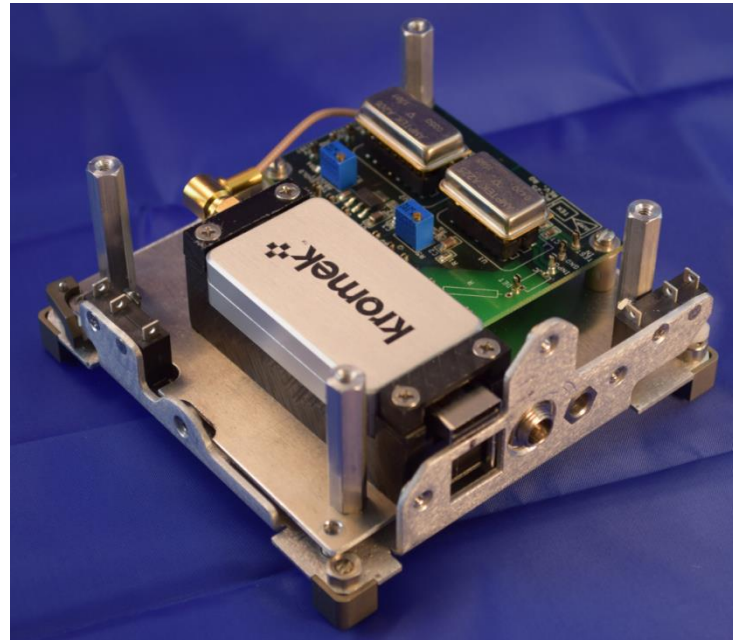


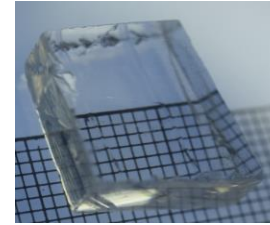
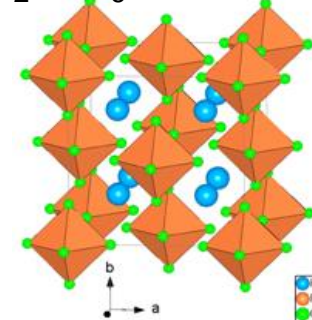
Z. Bell et al,
2015 SPIE Optics and Photonics,
International Conference on Hard X-Ray,
Gamma Ray and Neutron Detector
Physics



LiInSe₂ PROTOTYPE, CUBESAT 2U

- Collaboration funded by Vanderbilt – Discovery grant
- Space is sufficient to add a Raspberry Pi computer (Linux)
- LiInSe₂ operates in scintillator mode with a single APD readout. Plan to test it for neutron detection at LPL – University of Arizona





Cesium hafnium chloride: A high light yield, non-hygroscopic cubic crystal scintillator for gamma spectroscopy

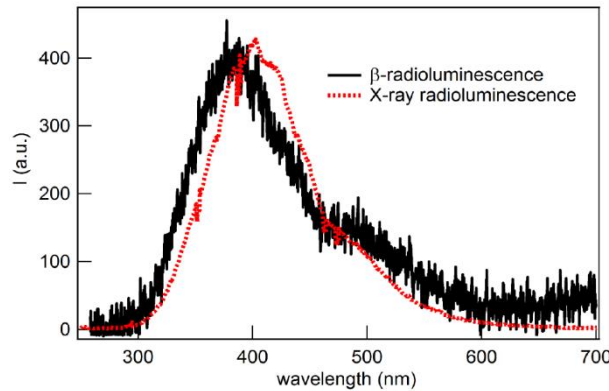
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Published in APL Vol 107, on the web at <http://dx.doi.org/10.1063/1.4932570>



$\lambda_{\max} = 400 \text{ nm}$

FIG. 1. Cs_2HfCl_6 radioluminescence acquired with a Sr-90 beta source is in good agreement with x-ray excited radioluminescence acquired with < 30 keV x-rays with Silicon CCD readout. Light yield is estimated to be 54,000 Ph/MeV by comparison to a standard BGO crystal. The slight shift between spectra is due to differences in spectral sensitivity between the two spectrometers used for the measurements.

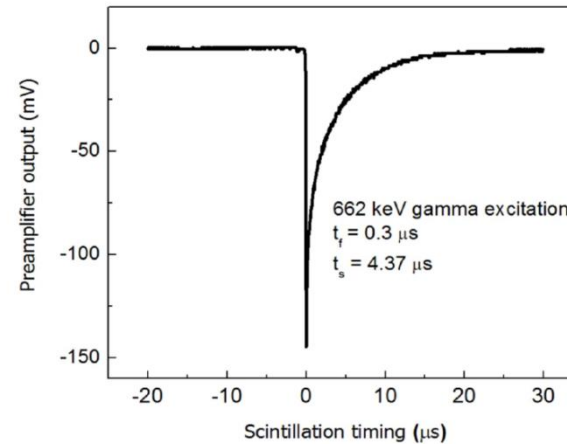
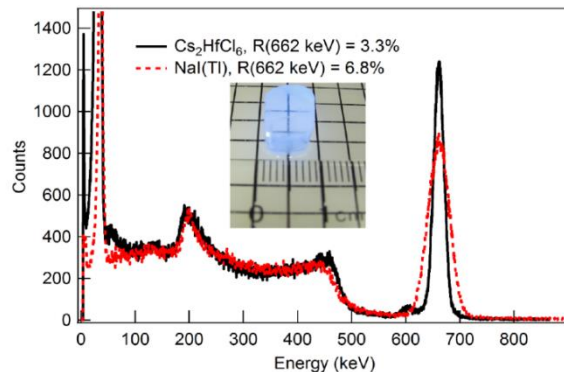


FIG. 2. Digital oscilloscope traces for timing measurements of the CHC scintillation response to 662 keV gammas. The decay time has a fast component (5% of energy) of 0.3 μs and a slow component (95% of energy) of 4.37 μs . The size of the CHC sample was 8.5 mm^3 .

principal
decay time is
4.37 μs



FWHM=3.3%
@ 662 keV

LY of 54,000 ph/MeV
with silicon CCD
photodetector

FIG. 3. Pulse height spectra acquired with a Cs-137 source of a Cs_2HfCl_6 crystal 0.65 cm^3 in size, compared to a standard 1" \times 1" NaI(Tl) crystal, both measured with 12 μs shaping time. The gamma light yield of Cs_2HfCl_6 , compared to NaI(Tl) is estimated to be 33,000 Photons/MeV, with super bialkali PMT readout. The small peak in the CHC spectrum at 607 keV is due to the escape of $K\alpha$ fluorescence x-ray, typical of small volume detectors. The inset is a photo of the Cs_2HfCl_6 crystal on a paper background (0.5 cm squares), under 254 nm excitation, exhibiting its characteristic blue emission.

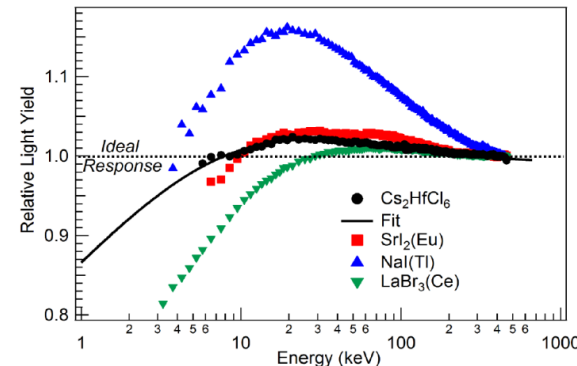


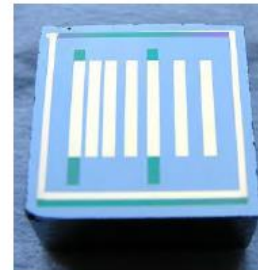
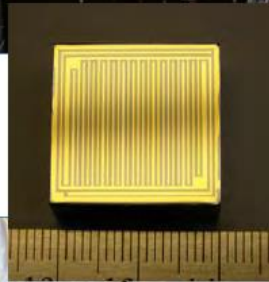
FIG. 4. Experimental electron light yield nonproportionality curves for Cs_2HfCl_6 , $\text{SrI}_2(\text{Eu})$, NaI(Tl), and $\text{LaBr}_3(\text{Ce})$. The ideal response would have no energy dependence of light yield. A fit to the CHC data indicates better proportionality than any previously reported scintillator.

non-proportionality-
limited
resolution
of 1.37% @ 662 keV

DETECTOR FABRICATION

Equipment:

diamond cutting wheel
diamond wire saws
8 in. Lapping Machine
e-beam evaporator
sputtering system
parlyene coating system
Photolitography system





That's all Folks!